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The influence of (n - n')-mixing processes in $He^*(n) + He(1s^2)$ collisions on $He^*(n)$ atoms' populations in weakly ionized helium plasmas

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Abstract

The results of semi-classical calculations of rate coefficients of (n-n')-mixing processes due to collisions of Rydberg atoms $\text{He}^*(n)$ with $\text{He}(1s^2)$ atoms are presented. It is assumed that these processes are caused by the resonant energy exchange within the electron component of $\text{He}^*(n)$ + He collision system. The method is realized through the numerical simulation of the (n-n')-mixing processes, and is applied for calculations of the corresponding rate coefficients. The calculations are performed for the principal quantum numbers n,n' in ranges $4 \le n < n' \le 10$, and the atom and electron temperatures, T_a, T_e , in domains $5000 \, \text{K} \le T_a \le T_e \le 20\,000 \, \text{K}$. It is shown that the (n-n')-mixing processes can significantly influence the populations of Rydberg atoms in non-equilibrium weakly ionized helium plasmas with ionization degree $\sim 10^{-4}$. Therefore, these processes have to be included in the appropriate models of such plasmas. © 2007 Elsevier Ltd. All rights reserved.

Keywords: Atomic and molecular collisions; Rydberg states

1. Introduction

Several existing models of collisional-radiative recombination [1–9] showed that in the weakly ionized plasmas, with ionization degree $\sim 10^{-4}$, the electron transfer from the lowest atomic states to the continuum and in the opposite direction is dominantly caused by the cascade mechanism. These processes of excitation/de-excitation are very important for populations in the lower part of the Rydberg's block of states. The reason for this is the distribution function of the excited atom states' populations in weakly ionized plasmas, which shows a distinct minimum ("bottleneck") in the lower part of Rydberg's block of states. For example, paper [10] compared LTE (local thermodynamical equilibrium) and non-LTE distribution functions, corresponding to different layers of the solar photosphere [5,6]. It was found that both functions had distinct minimum in the domain of principal quantum numbers n = 5 - 6. Such behavior is illustrated in Fig. 1, showing LTE distribution functions for hydrogen plasmas at the temperatures T = 5000 and 6000 K.

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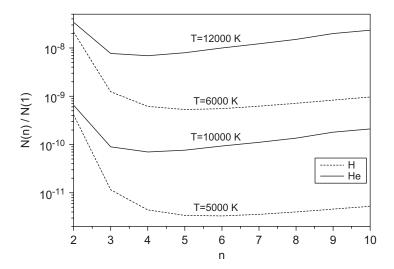


Fig. 1. The distribution function of excited atom population.

Similar behavior of the distribution functions for excited atomic states in the case of weakly ionized helium plasmas, for the temperatures $T = 10\,000$ and $12\,000$ K typical for photospheres of some helium rich white dwarfs [11], is illustrated in the same figure. The experimental data for weakly ionized plasma of non-equilibrium helium arc [12], show that the distribution function has a similar shape, even in the case when the electron temperature is 4 times higher than atomic temperature. It can be concluded that the kinetics of weakly ionized hydrogen and helium plasmas strongly depends on excitation and deexcitation processes in the region $4 \lesssim n \lesssim 10$, where the "bottleneck" of distribution function occurs.

The models of collisional-radiative recombination, used to estimate the influence of non-elastic processes which populate the lower part of Rydberg's block of states, included the radiation processes and the excitation/de-excitation processes caused by electron—atom collisions. The latter was assumed to be dominant for plasmas with the ionization degree $\sim 10^{-4}$. However, the distribution of excited atom states' populations experimentally obtained in [12], showed a significant influence of (n-n')-mixing processes in $He^*(n) + He(1s^2)$ collisions. It was found that the distribution function was similar to Boltzmann with the temperature about $11\,000\,\mathrm{K}$, close to the mid-value between $T_a \cong 4500\,\mathrm{K}$ and $T_e \cong 18\,000\,\mathrm{K}$. Consequently, the mechanism of the resonant energy exchange within the electron component of the atom-Rydberg atom collision systems (resonant mechanism) has been introduced in [13] to explain such a behavior. In cited paper and in [14,15] the resonant mechanism was tested only for (n-n')-mixing processes in $H^*(n) + H(1s)$ collisions, and the results obtained had only a qualitative character.

The resonant mechanism has been successfully applied to chemi-ionization processes in alkali Rydberg's atoms collisions with their parents' ground state [16], and further for alkali atoms' collisions [17–26]. In the case of hydrogen and helium atoms, the resonant mechanism was used to analyze the chemi-ionization and their inverse chemi-recombination processes [27–33]. It was shown that in weakly ionized hydrogen and helium plasmas the chemi-ionization/recombination processes caused by the resonant mechanism can be dominant in populating the Rydberg's states. This mechanism was also investigated in the case of (n-n')-mixing processes in $H^*(n) + H$ collisions populating the states within the Rydberg's block [10] (see also [14,15]). In [10] the corresponding rate coefficients were determined as functions of n and T_a . It was shown that in hydrogen plasma with the ionization degree $\sim 10^{-4}$, and $4 \le n \le 10$, the efficiency of these processes is comparable or even higher than the concurrent excitation/de-excitation processes caused by electron—atom collisions. This was confirmed later for weakly ionized hydrogen plasma of solar photosphere in [34].

The efficiency of the resonant mechanism for chemi-ionization/recombination processes in weakly ionized hydrogen and helium plasmas [29,30] suggested that one should expect similar situation in the case of (n - n')-mixing processes. The semi-classical method based on this mechanism, developed in [10], was used in this

paper and applied for the rate coefficients' calculations of excitation processes

$$He^*(n) + He \rightarrow \begin{cases} He^*(n' = n + p) + He, \\ He + He^*(n' = n + p) \end{cases}$$
 (1)

and their inverse deexcitation processes

$$He^*(n) + He \rightarrow \begin{cases} He^*(n' = n - p) + He, \\ He + He^*(n' = n - p) \end{cases}$$
 (2)

with $p \ge 1$. The processes are characterized by the excitation and deexcitation rate coefficients $K_{n;n+p}(T_a)$ and $K_{n;n-p}(T_a)$, respectively, where T_a is the atom temperature. They represent the average characteristics of transitions from all individual states with a given n to all states with given n+p or n-p, and reflect the processes' influence on the plasma kinetics. The calculations are performed for the parameters' domains $4 \le n \le 10$, $1 \le p \le 5$ and $4000 \text{ K} \le T_a \le T_e \le 2000 \text{ K}$. The efficiency of processes (1) and (2) is compared to the concurrent electron-excited atom collisional excitation/de-excitation processes

$$He^*(n) + e \to He^*(n \pm p) + e',$$
 (3)

where e and e' denote free electron in the initial and final state, for given n, p and electron $T_{\rm e}$.

Beside of processes (3) in the plasma occur the other concurrent processes, including the radiative processes caused by the interaction of excited atoms with the electromagnetic radiation. However, just the processes (3) are taken as referent ones since they are always included in the modeling of weakly ionized plasma disregarding to other concurrent processes. Consequently, in order to demonstrate the necessity of including of the processes (1) and (2) in the models of weakly ionized plasmas, it is enough to show that their efficiency is comparable with the efficiency of processes (3).

2. Theoretical remarks

Resonant mechanism: The resonant mechanism of non-elastic (n - n')-mixing and chemi-ionization/recombination processes in $A^*(n) + A$ collisions was described in details in previous papers [10,14,17,27,28]. In considered cases A was an atom in the ground state with one or two s-electrons out of completely filled shells. The mechanism, briefly described here, is illustrated by Figs. 2a and b.

Fig. 2a shows that the resonant mechanism region is defined by the inequality

$$R \ll r_n$$
, (4)

where R is the internuclear distance and $r_n \sim n^2$ the characteristic radius of Rydberg's atom $A^*(n)$. In this region a perturbation operator takes the dipole part of the interaction between the outer electron e_n and the sub-system $A^+ + A$.

The electronic state of the $A^+ + A$ sub-system is described by the wave functions of the adiabatic ground ($|1;R\rangle$) and first excited ($|2;R\rangle$) electronic states of the molecular ion A_2^+ . The potential curves of these states are denoted with $U_1(R)$ and $U_2(R)$, and schematically shown in Fig. 2b. The impact energy of the collisional system, determined in the center of mass' system, is denoted with E. Fig. 2b illustrates that the resonant mechanism takes into account the transitions $|n\rangle \to |n+p\rangle$, $|n\rangle \to |\vec{k}\rangle$ and $|\vec{k}\rangle \to |n\rangle$, $|n\rangle \to |n-p\rangle$ of the outer electron, which occur simultaneously with the transitions $|2;R\rangle \to |1;R\rangle$ and $|1;R\rangle \to |2;R\rangle$ in the subsystem $A^+ + A$.

The main cause of the resonant character of described mechanism is the importance of transitions in the close vicinity of the corresponding resonant distances. In the case of (n - n')-mixing, the resonant distances $R_{n,n\pm p}$ are found as roots of the equation

$$U_{12}(R) \equiv U_2(R) - U_1(R) = |\varepsilon_{n \pm p} - \varepsilon_n|,$$
 (5)

where ε_n and $\varepsilon_{n\pm p}$ denote electron energies in bound states with principal quantum numbers n and $n\pm p$. Note that for the chemi-ionization processes the resonant distances $R_{n;k}$ are found from (5) by replacing $\varepsilon_{n\pm p}$ with the energy of the outer electron in free state, ε_k .

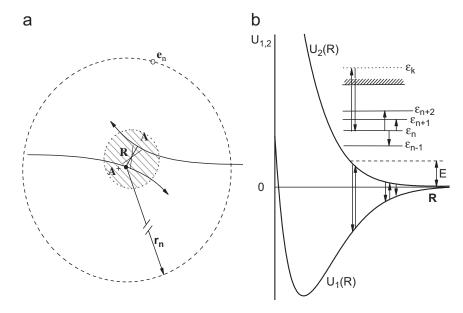


Fig. 2. (a) Schematic illustration of $He^*(n) + He$ collision (the region of the inter-nuclear distance R where the outer electron is collectivized is shaded). (b) Schematic illustration of the simultaneous resonant transitions of the outer electron from the initial bound to the final free state and the sub-system $He^+ + He$ from initial excited to the final ground electronic state.

Table 1 $R_{n,n+p}/a_0$ values calculated from (5) for n' = n + p

n	p .								
	1	2	3	4	5				
4	4.907	4.556	4.400	4.312	4.257				
5	5.398	5.019	4.843	4.741	4.675				
6	5.802	5.403	5.211	5.097	5.022				
7	6.146	5.731	5.527	5.403	5.321				
8	6.446	6.018	5.803	5.672	5.582				
9	6.711	6.272	6.050	5.911	5.816				
10	6.949	6.502	6.272	6.128	6.028				

The described resonant mechanism for (n-n')-mixing processes (1) and (2) is applied in helium case: $|1;R\rangle$ and $|2;R\rangle$ are $X^2\Sigma_u^+$ - and $A^2\Sigma_g^+$ -states of the molecular ion He_2^+ . The adiabatic potential curves $U_1(R)$ and $U_2(R)$ are taken from [35,36]. The resonant distances $R_{n;n\pm p}$ are determined from Eq. (5) with these potential curves. The values of $R_{n;n+p}$ for $4 \le n \le 10$ and $1 \le p \le 5$ are presented in Table 1.

The rate coefficients' calculations require the quantity $D_{12}(R) = |\langle 1; R|\mathbf{D}|2; R\rangle|$, where **D** is the operator of the dipole momentum of the molecular ion He_2^+ . Table 1 shows that for all values of $R_{n;n+p}$ the inequality $R \gg a_0$ holds (a_0 being the atomic unit of length), and for D_{12} the approximation from [10] can be used

$$D_{12}(R) = \frac{e \cdot R}{2},\tag{6}$$

where e is the absolute value of electron charge.

The method of calculations: $\text{He}^*(n \ge 4) + \text{He}$ and $\text{H}^*(n \ge 4) + \text{H}$ collision systems have similar behavior in $R \ge a_0$ region, which has been used for the chemi-ionization processes' calculations [30,33]. On the other hand, the resonant mechanism generates transitions between Rydberg's states of outer electron that are allowed by selective rules. This was used to apply a semi-classical theory for the chemi-ionization processes in the cases of hydrogen [28,29] and helium [30,33].

The (n-n')-mixing processes in $He^*(n)$ + He collisions are considered to be a result of the same resonant mechanism as in previous cases. The similarities between hydrogen and helium collisional systems made it possible to apply a semi-classical theory from [10] for processes (1) and (2). It means that the rate coefficients for these processes can be determined by the same procedure used in [10], but with the molecular ion He_2^+ characteristics instead. Consequently, the rate coefficients $K_{n;n+p}(T_a)$ for excitation processes (1) are determined directly, and the coefficients $K_{n;n-p}(T_a)$ for deexcitation processes (2) by using thermodynamical balance principle.

Similar to [10], the rate coefficients $K_{n,n+p}(T_a)$ are calculated under assumptions that:

- the internuclear motion occurs with the probability $\frac{1}{2}$ in the potential $U_2(R)$ (molecular state $|2;R\rangle$), and can be described by the classical trajectory determined by the impact parameter and energy E;
- the transitions $|2;R\rangle|n\rangle \rightarrow |1;R\rangle|n+p\rangle$ for p=1,2,... happen continually along the trajectory, and are characterized by their rates;
- ullet the atomic component has a Maxwell distribution described by the impact energies E with a given temperature $T_{\rm a}$.

Applying the method from [10], the excitation rate coefficients $K_{n;n+p}(T_a)$ are given as

$$K_{n;n+p}(T_{\rm a}) = \frac{2\pi}{3\sqrt{3}} \frac{(ea_0)^2}{\hbar} \cdot n^{-5} \cdot g_{n;n+p} \int_{R_{\rm min}(n,n+p)}^{R_{\rm max}(n,n+p)} X(R) \cdot \exp\left[-\frac{U_2(R)}{kT_{\rm a}}\right] \frac{R^4 \cdot dR}{a_0^5},\tag{7}$$

where $g_{n,n+p}$ is Gaunt's factor for $|n\rangle \to |n+p\rangle$ transition, determined in [37]. The lower and upper bounds $R_{\min}(n,n+p)$ and $R_{\max}(n,n+p)$ are found as roots of the equations

$$U_{12}(R = R_{\min}) = \varepsilon_{n+p+1-\Delta ap} - \varepsilon_n, \quad U_{12}(R = R_{\max}) = \varepsilon_{n+p-\Delta p} - \varepsilon_n$$
(8)

with $\varepsilon_n = -Ry/n^2$ and $\Delta p = 0.380$. A factor X(R) is given by

$$X(R) = \frac{\Gamma(\frac{3}{2}, -U_1(R)/kT_a)}{\Gamma(\frac{3}{2})},\tag{9}$$

and $\Gamma(x, y)$ and $\Gamma(x)$ are incomplete and complete gamma functions [38].

The deexcitation rate coefficients $K_{n-p;n}(T_a)$ are calculated from the thermodynamic balance equation in the case of helium plasma in the state of thermodynamical equilibrium with $T = T_a$, namely

$$K_{n-p;n}(T=T_a)N_{eq}(n-p)N_{eq}(1) = K_{n;n-p}(T=T_a)N_{eq}(n)N_{eq}(1),$$
(10)

where $N_{\rm eq}(1)$, $N_{\rm eq}(n-p)$ and $N_{\rm eq}(n)$ are the corresponding equilibrium densities of the atoms He(1s²), He*(n-p) and He*(n), respectively. These densities satisfy equations

$$N_{\text{eq}}(n-p) = N_{\text{eq}}(1) \cdot 4(n-p)^2 \exp\left(-\frac{I_{\text{He}} + \varepsilon_{n-p}}{kT_{\text{a}}}\right),$$

$$N_{\text{eq}}(n) = N_{\text{eq}}(1) \cdot 4n^2 \exp\left(-\frac{I_{\text{He}} + \varepsilon_n}{kT_{\text{a}}}\right),$$
(11)

where I_{He} is the ionization potential of $\text{He}(1s^2)$ atom. Consequently, the relation which links the excitation and deexcitation rate coefficients is

$$K_{n;n-p}(T_{\mathbf{a}}) = K_{n-p;n}(T_{\mathbf{a}}) \cdot \frac{(n-p)^2}{n^2} \cdot \exp\left(\frac{\varepsilon_{n-p;n}}{kT_{\mathbf{a}}}\right),\tag{12}$$

where $\varepsilon_{n-p,n} = \varepsilon_n - \varepsilon_{n-p}$. The usage of the thermodynamical balance principle understands that in the equilibrium helium plasma with $T = T_a$ Boltzmann's distribution of helium excited atom populations has its full physical sense. One can directly confirm that this is valid for $T_a \gtrsim 3000$ K. Because of that the described procedure can be certainly applied in the temperature region considered here.

3. Results and discussion

Eqs. (7)–(9) are used for rate coefficients $K_{n;n+p}(T_a)$ calculations. The results for $4 \le n \le 10$, $1 \le p \le 5$ and $4000 \, \mathrm{K} \le T_a \le 20\,000 \, \mathrm{K}$ are presented in Table 2. The Table shows a monotonous decrease of $K_{n;n+p}(T_a)$ with the increases of n and p, and a very slow increase as T_a increases. It should be noted that the differences between $K_{n;n+p}$ values for hydrogen from [10] and these for helium are caused by differences between the resonant distances $R_{n;n+p}$. The deexcitation rate coefficients $K_{n;n-p}(T_a)$ for $4 < n \le 10$ and $4 \le n-p$ are calculated according to Eq. (12) and using the values for excitation rate coefficient from Table 2.

The relative efficiency of (n - n')-mixing processes (1) and (2) in comparison with electron-atom collision processes (3) is characterized by the quantity $F_{n,n\pm p}(T_a, T_e)$

$$F_{n;n\pm p}(T_{\rm a}, T_{\rm e}) = \frac{K_{n;n\pm p}(T_{\rm a})N(n)N(1)}{\alpha_{n;n\pm p}(T_{\rm e})N(n)N_{\rm e}} = \frac{K_{n;n\pm p}(T_{\rm a})}{\alpha_{n;n\pm p}(T_{\rm e})} \cdot \eta_{\rm ea},\tag{13}$$

Table 2 Excitation rate coefficients $K_{n,n+p}(T_a)$ (10⁻⁹ cm³ s⁻¹)

n	p	$T_{\rm a}~(10^3~{ m K})$								
		4.0	4.5	5.0	6.0	8.0	10.0	12.0	16.0	20.0
4	1	1.005	1.080	1.143	1.243	1.378	1.462	1.520	1.594	1.638
	2	0.195	0.221	0.244	0.283	0.338	0.374	0.400	0.434	0.455
	3	0.072	0.084	0.095	0.114	0.142	0.161	0.175	0.194	0.206
	4	0.035	0.041	0.048	0.058	0.075	0.087	0.095	0.107	0.114
	5	0.020	0.024	0.028	0.035	0.045	0.053	0.058	0.066	0.071
5	1	0.682	0.707	0.728	0.760	0.800	0.824	0.841	0.861	0.873
	2	0.170	0.182	0.191	0.207	0.227	0.239	0.248	0.258	0.265
	3	0.072	0.078	0.083	0.092	0.103	0.111	0.116	0.122	0.126
	4	0.038	0.042	0.045	0.050	0.058	0.062	0.066	0.070	0.072
	5	0.023	0.025	0.027	0.031	0.036	0.039	0.041	0.044	0.046
6	1	0.432	0.441	0.448	0.459	0.473	0.481	0.486	0.493	0.497
	2	0.124	0.128	0.132	0.138	0.146	0.151	0.154	0.157	0.160
	3	0.056	0.059	0.062	0.065	0.070	0.073	0.075	0.077	0.079
	4	0.031	0.033	0.035	0.037	0.040	0.042	0.044	0.045	0.046
	5	0.019	0.021	0.022	0.024	0.026	0.027	0.028	0.029	0.030
7	1	0.276	0.279	0.282	0.286	0.292	0.295	0.297	0.299	0.301
	2	0.086	0.088	0.089	0.092	0.095	0.097	0.098	0.100	0.101
	3	0.041	0.042	0.044	0.045	0.047	0.049	0.049	0.050	0.051
	4	0.024	0.025	0.025	0.027	0.028	0.029	0.029	0.030	0.031
	5	0.015	0.016	0.016	0.017	0.018	0.019	0.019	0.020	0.020
8	1	0.181	0.183	0.184	0.186	0.188	0.189	0.190	0.191	0.192
	2	0.059	0.060	0.061	0.062	0.064	0.065	0.065	0.066	0.066
	3	0.030	0.030	0.031	0.032	0.033	0.033	0.034	0.034	0.034
	4	0.018	0.018	0.018	0.019	0.020	0.020	0.020	0.021	0.021
	5	0.011	0.012	0.012	0.013	0.013	0.013	0.014	0.014	0.014
9	1	0.122	0.123	0.124	0.125	0.126	0.126	0.127	0.127	0.128
	2	0.042	0.042	0.043	0.043	0.044	0.044	0.045	0.045	0.045
	3	0.021	0.022	0.022	0.022	0.023	0.023	0.023	0.024	0.024
	4	0.013	0.013	0.013	0.014	0.014	0.014	0.014	0.015	0.015
	5	0.009	0.009	0.009	0.009	0.009	0.010	0.010	0.010	0.010
10	1	0.085	0.086	0.086	0.086	0.087	0.087	0.087	0.088	0.088
	2	0.030	0.030	0.031	0.031	0.031	0.031	0.032	0.032	0.032
	3	0.016	0.016	0.016	0.016	0.017	0.017	0.017	0.017	0.017
	4	0.010	0.010	0.010	0.010	0.010	0.010	0.010	0.011	0.011
	5	0.006	0.007	0.007	0.007	0.007	0.007	0.007	0.007	0.007

$$\eta_{\rm ea} = \frac{N(1)}{N_c},\tag{14}$$

where $\alpha_{n,n\pm p}(T_e)$ are the rate coefficients for the electron-atom excitation processes (3), and N_e is the free electron density. In the case of weakly ionized helium plasma the parameter η is closed to (ionization degree)⁻¹. The values for $\alpha_{n,n\pm p}(T_e)$ are obtained from [39,40].

The quantities $F_{n;n\pm p}(T_a, T_e)$ as functions of n for given p, T_a and T_e are illustrated in Figs. 3–9. They show that the efficiency of (n-n')-mixing processes in He*(n) + He collisions, in domains $4 \le n \le 8$ and $n+1 \le n' \le n+5$, is higher or at least comparable to that of electron–atom processes (3), for all considered T_a and T_e .

The region $4 \le n \le 6$, where the minimum of helium excited atom populations' distribution occurs, is particularly important (see Fig. 1). In this domain the efficiency of processes (1) and (2) is significantly higher or very close to that of electron—atom processes (3). Such efficiency can explain the shape of distribution function, experimentally obtained in [12], and discussed in Section 1.

Regarding to the presented method accuracy we rely to the fact that (n - n')-mixing processes are caused by the same resonant mechanism as the above mentioned chemi-ionization processes in symmetric atom-Rydberg atom collisions. Namely, these processes are well studied experimentally and theoretically and it was found that there is very good agreement between experimental and theoretical results [17–26].

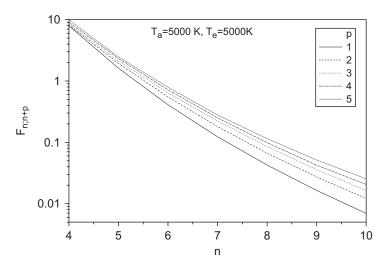


Fig. 3. Parameter $F_{n,n+p}$ for $T_a = 5000 \,\mathrm{K}$ and $T_e = 5000 \,\mathrm{K}$.

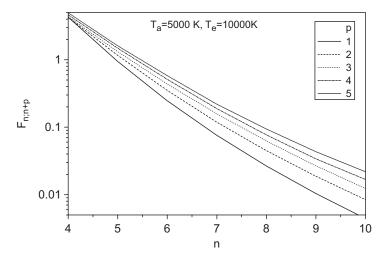


Fig. 4. Same as in Fig. 3, but for $T_a = 5000 \, \mathrm{K}$ and $T_e = 10000 \, \mathrm{K}$.

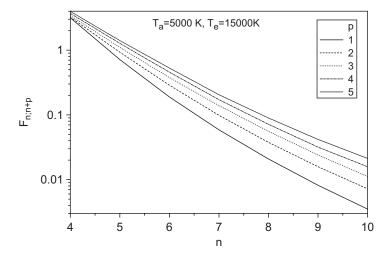


Fig. 5. Same as in Fig. 3, but for $T_a = 5000 \, \mathrm{K}$ and $T_e = 15000 \, \mathrm{K}$.

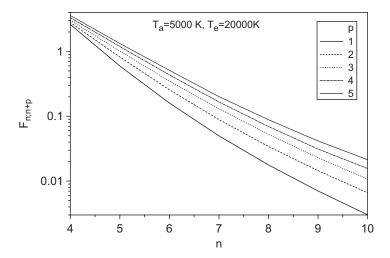


Fig. 6. Same as in Fig. 3, but for $T_{\rm a}=5000\,{\rm K}$ and $T_{\rm e}=20\,000\,{\rm K}$.

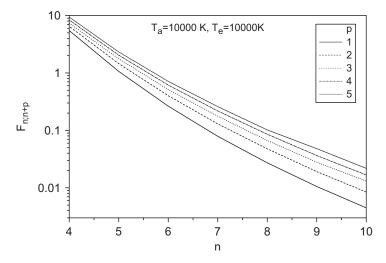


Fig. 7. Same as in Fig. 3, but for $T_{\rm a}=10\,000\,{\rm K}$ and $T_{\rm e}=10\,000\,{\rm K}$.

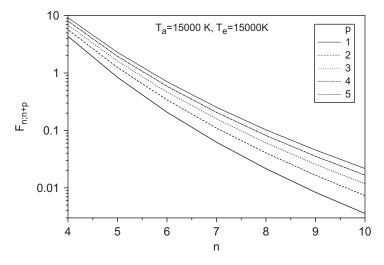


Fig. 8. Same as in Fig. 3, but for $T_a = 15000 \,\mathrm{K}$ and $T_e = 15000 \,\mathrm{K}$.

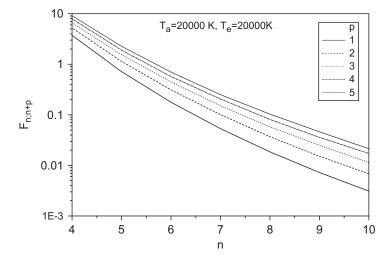


Fig. 9. Same as in Fig. 3, but for $T_a = 20\,000\,\mathrm{K}$ and $T_e = 20\,000\,\mathrm{K}$.

The results of this paper suggest that the influence of the processes (1) and (2) on populating the lower part of Rydberg block of states has to be taken into account. It was shown in [32] that the similar mechanism in the chemi-ionization processes in weakly ionized hydrogen plasma, in region $4 \le n \le 8$, significantly changed the excited atom populations in the entire Rydberg block of states, and hence the changes in atomic spectral line shapes [41]. Therefore, the role of (n - n')-mixing processes for the populations of the lower part of Rydberg block of states is of a special importance, the fact that has already been confirmed in the case of hydrogen [10].

The importance of chemi-ionization/recombination processes in weakly ionized helium plasmas has been confirmed earlier [30,33], particularly when they were included within a model of collision radiative recombination [9]. On the other hands, the results presented here suggest a similar role for (n - n')-mixing processes (1) and (2) in both helium and hydrogen plasmas. For example, these processes should influence helium spectral lines' shapes in cases of some helium rich white dwarfs atmospheres, in a research which already started [42]. They are also important for the kinetics of weakly ionized laboratory plasmas, which are always used in a research of gaseous discharges and inner-plasma collision processes [12,43–48]. To conclude, the (n - n')-mixing processes (1) and (2) have to be included in modeling of both laboratory and astrophysical weakly ionized helium plasmas.

References

- [1] Bates DR, Kingston AE, McWhirter RWP. Proc R Soc 1962;267:297.
- [2] Bates DR, Kingston AE. Proc R Soc 1964;279:32.
- [3] Biberman LM, Vorobjev VS, Yakubov IT. Teplofiz Vys Temp 1967;5:201.
- [4] Biberman LM, Vorobjev VS, Yakubov IT. Uspehi Fiziceskih Nauk 1979;128:233.
- [5] Vernazza J, Avrett E, Loser R. Astrophys J Suppl Ser 1981;45:635.
- [6] Maltby P, Avertt EH, Carlsson M, Kjeldseth-Moe O, Kurucz RL, Loeser R. Astrophs J 1986;306:284.
- [7] Biberman LM, Vorobjev VS, Yakubov IT. Kinetics of nonequilibrium low-temperature plasma. New York: Plenum Press; 1987.
- [8] Pert GJ. J Phys B 1990;23:619.
- [9] Djurić Z, Mihajlov AA. JQSRT 2001;70:285.
- [10] Mihajlov AA, Ignjatović LM, Djurić Z, Ljepojević NN. J Phys B 2004;37:4493.
- [11] Koester D. Astrophys J Suppl Ser 1980;39:401.
- [12] Aleksandrov VY, Gurevich DB, Podmoshenskii IV. Opt Spektrosk 1969;36-41.
- [13] Smirnov VM, Mihajlov AA. Opt Spektrosk 1971;30:984.
- [14] Janev RK, Mihajlov AA. Phys Rev A 1979;20:1890.
- [15] Mihajlov AA. In: Proceedings of international conference on plasma physics. G 1982.
- [16] Devdariani AZ, Klyucharev AN, Lazarenko AB, Sheverev VA. Pis'ma Zh Tekh Fiz 1978;4:1013.
- [17] Mihajlov AA, Janev RK. J Phys B 1981;14:1639.
- [18] Weiner J, Polak-Dingels P. J Chem Phys 1981;74:508.
- [19] Weiner J, Masnou-Seeuws F, Giusti-Suzor A. Adv A Mol Opt Phys 1989;26:209.
- [20] Klucharev AN, Vujnović V. Phys Rep 1990;185:55.
- [21] Bezuglov NN, Borodin VM, Klyucharev AN, Fuso F, Allegrini M. Opt Spectrosc 1997;83:338.
- [22] Bezuglov NN, Borodin VM, Kazanskiy AK, Klyucharev AN, Matveev AA, Orlovskiy KV. Opt Spectrosc 2001;91:19.
- [23] Ryabtsev II, Tretyakov DB, Beterov II, Bezuglov NN, Miculis K, Ekers A. J Phys B 2005;38:S17.
- [24] Beterov I, Tretyakov D, Ryabtsev I, Bezuglov N, Miculis K, Ekers A, et al. J Phys B 2005;38:4349.
- [25] Miculis K, Beterov I, Bezuglov N, Ryabtsev I, Tretyakov D, Ekers A, et al. J Phys B 2005;38:1811.
- [26] Ignjatović LM, Mihajlov AA. Phys Rev A 2005;72:022715.
- [27] Janev RK, Mihajlov AA. Phys Rev A 1980;21:819.
- [28] Mihajlov AA, Ljepojević NN, Dimitrijević MS. J Phys B 1992;25:5121.
- [29] Mihajlov AA, Dimitrijević MS, Djurić Z. Phys Scr 1996;53:159.
- [30] Mihajlov AA, Dimitrijević MS, Djurić Z, Ljepojević NN. Phys Scr 1997;56:631.
- [31] Mihajlov AA, Ignjatovic LM, Vasilijević MM, Dimitrijević MS. Astron Astrophys 1997;324:1206.
- [32] Mihajlov AA, Jevremović D, Hauschildt P, Dimitrijević MS, Ignjatović LM, Alard F. Astron Astrophys 2003;403:787.
- [33] Mihajlov AA, Ignjatović LM, Dimitrijević MS, Djurić Z. The Astrophys J Suppl Ser 2003;147:369.
- [34] Mihajlov AA, Ignjatović LM, Dimitrijević MS. Astron Astrophys 2005;437:1023.
- [35] Gupta BM, Madsen FA. J Chem Phys 1967;47:48.
- [36] Metropoulos A, Li Y, Hirsch G, Buenker RJ. Chem Phys Lett 1992;198:266.
- [37] Johnson LC. Astrophys J 1972;174:227.
- [38] Abramowitz M, Stegun IA. Handbook of mathematical functions. New York: Dover; 1965.
- [39] Fujimoto T. Institute of plasma physics. Japan: Nagoya; 1978.
- [40] Janev RK, Langer WD, Evans Jr K, Post Jr DE. Elementary processes in hydrogen-helium plasmas. Berlin: Springer; 1987.
- [41] Mihajlov AA, Jevremović D, Hauschildt P, Dimitrijević MS, Ignjatović LM, Alard F. Astron Astrophys 2007;471:671.
- [42] Beauchamp A, Wesemael F, Bergeron P. Astrophys J Suppl Ser 1997;108:559.
- [43] Aleksandrov VY, Gurevich DB, Mihajlov AA, Podmoshenskii IV. Opt Spektrosk 1974;37:855.
- [44] Pendleton WR, Larsson M, Mannfors B. Phys Rev A 1985;28:3223.
- [45] Kiselevskii LI, Mazurenko SL, Makarevich AN. In: Proceedings of ICPIA-18, vol. 3. Swansea: 1987. p. 570.
- [46] Bayarkhuu C, Kiselevskii LI, Mazurenko SL, Makarevich AN, Solov'yanchik DA. J Appl Spectros 1989;51:758.
- [47] Solov'yanchik DA. J Appl Spectros 1993;59:340.
- [48] Denkelmann R, Maurmann S, Lokajczyká T, Drepper P, Kunze HJ. J Phys B: At Mol Opt Phys 1999;32:4635.